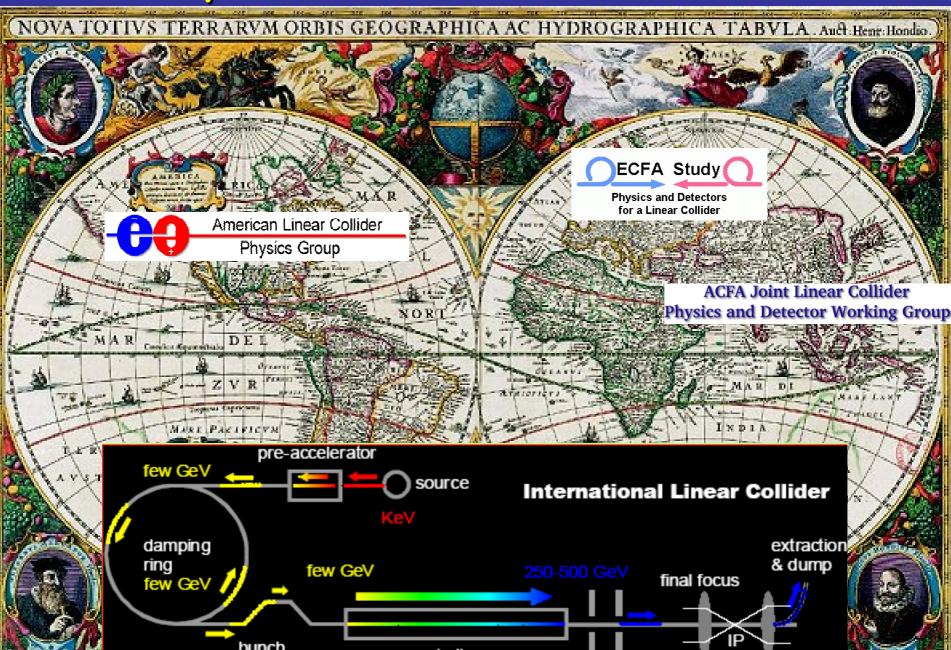
ILC: Physics and Politics issues



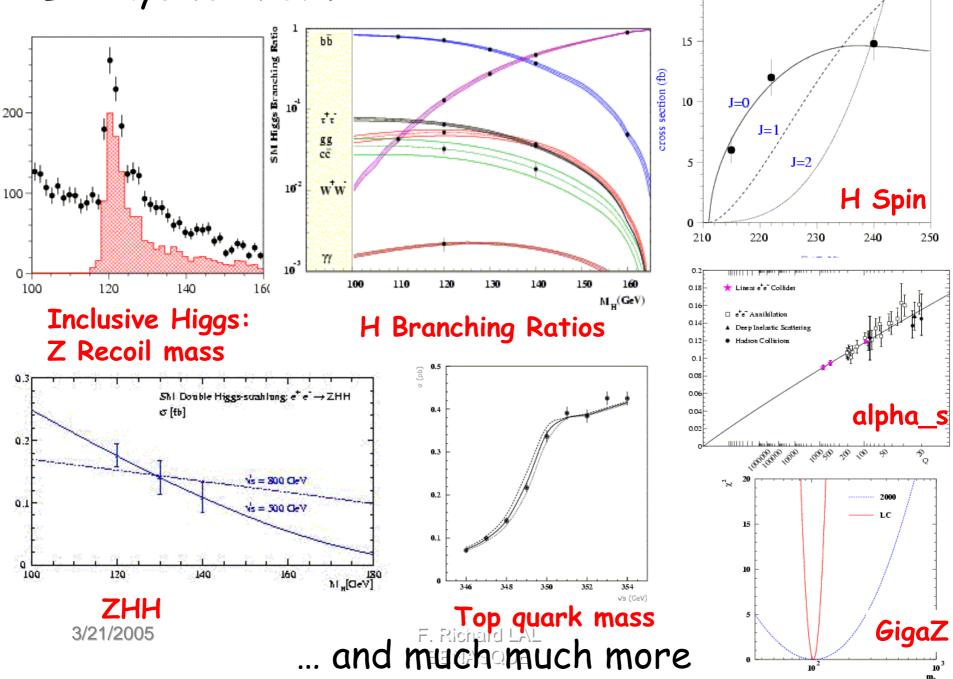
Purpose of this presentation

- Mainly on physics issues, focused on the SUSY scenario
- Few aspects of the detector and the machine
- Overview on the strategy

Physics at ILC

- 3 topics:
- EWSB: Higgs + ttbar, requires ILC up to 1 TeV with high L, very well covered by past studies (1 TeV needed on the S.I. scenario)
- SUSY: best scenario so far but not sure to 'exist' that is to be seen at LHC/ILC
- PM which allow in some cases to reach mass scales well beyond LHC/LC

ILC Physics in SM



SCOPE of the ILC

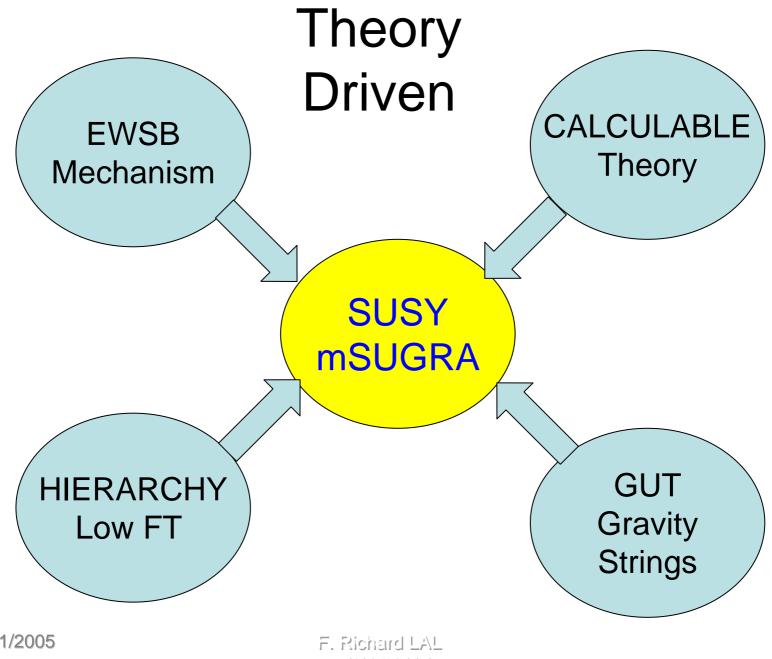
- Set by the physics
- 500 GeV phase to study Higgs+top+PM cumulating 500 fb-1 in ~6 y
- ~1 TeV phase to reach ~1 ab-1 needed e.g. for Zhh, to extend the reach of discoveries, to cover the SI scenario (WL-WL) if needed
- Options:

e+ polarized needed for GigaZ

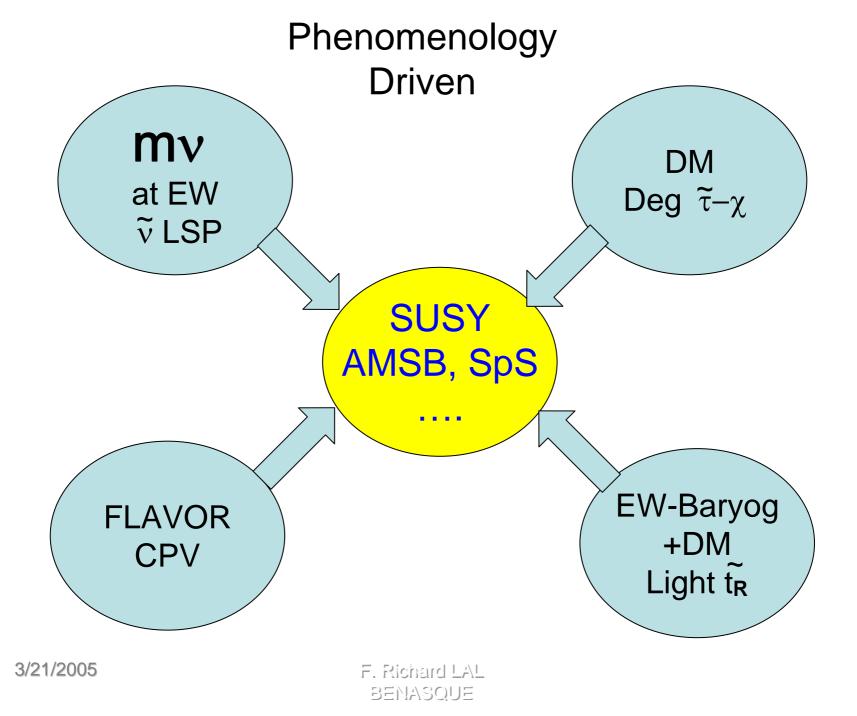
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SUSY

- At present the only ~ satisfactory extension of SM but no solid predictions on masses
- LHC/LC benchmarks studied within mSUGRA
- Several issues related to flavor and CPV, require different schemes 'protecting flavor' Gauge/Gaugino –mediated, SpS, AMSB
- DM plays a central role (also the gravitino pb)-> connection to cosmology
- -> Various schemes proposed either 'top-down' or bottom-up' or some mixture



BENASQUE

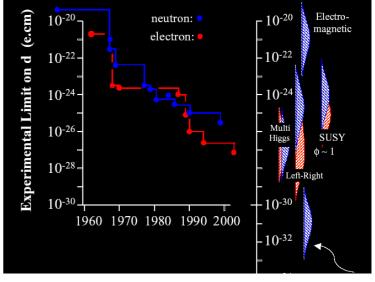


CPV and SUSY

E Richard AL

Two usual ways to satisfy flavor and CPV issues:

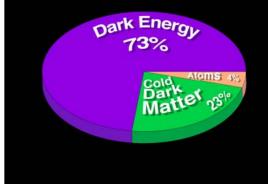
- Complete decoupling from flavor like AMSB
 Very heavy gravitinos (no reheating issues)
 Light sleptons (g-2 OK, v sector without see-saw)
- Send all scalars to very high masses like SpS
 Keeps the 'good part' of SUSY
 GUT, neutralino LSP
 CPV signals expected
 Highly FT but Nature also !



SpS	AMSB	ν _R
Higgsino	Wino	$\tilde{\mathbf{v}}$
No	Yes	Yes
Yes	No	?
170 GeV	Fat Higgs	
Not certain	Not certain	Confusing
(Gluino)	(LSP 2 TeV)	
See-saw ?	Dirac	Majorana
	Higgsino No Yes 170 GeV Not certain (Gluino)	Higgsino Wino No Yes Yes No 170 GeV Fat Higgs Not certain (Gluino) Kot certain

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As an example: the DM issue



- DM constrains severely SUSY
- There are 3 types of scenarios:
- The LSP is a Bino, hard to annihilate

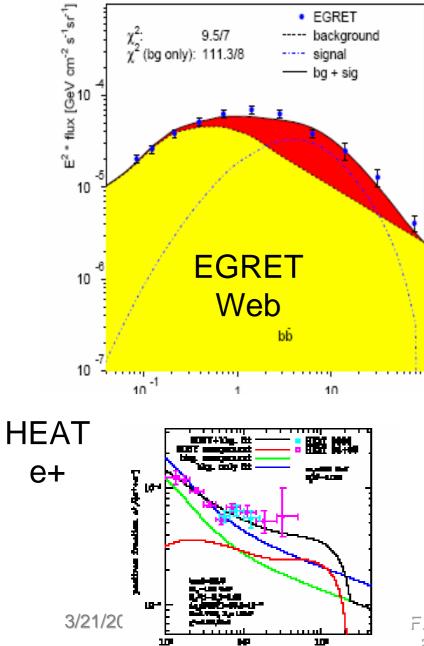
-> either very light sleptons or sleptons degenerate in mass (< 10 GeV) with LSP

- The LSP has a Higgsino component (low μ) sleptons very heavy SpS
- The LSP is a Wino AMSB. To saturate WMAP, the LSP has to be 2 TeV

Has DM being discovered ?!

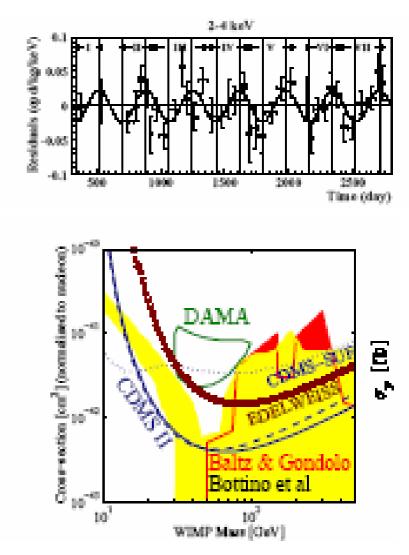
- Several indications
 Direct detection > 6 sd in DAMA
 Indirect excess reported in EGRET, HEAT
- Cannot be certain that they have common origin (modeling, reduced observables)
- Lesson (P. Gondolo, G. Kane): we will need LHC/ILC to be sure of the interpretation and generate a proper cosmological modeling
- We will need ILC to be precise enough to match Planck precision and make sure that there is a significant coincidence

Astro-ph/0408272



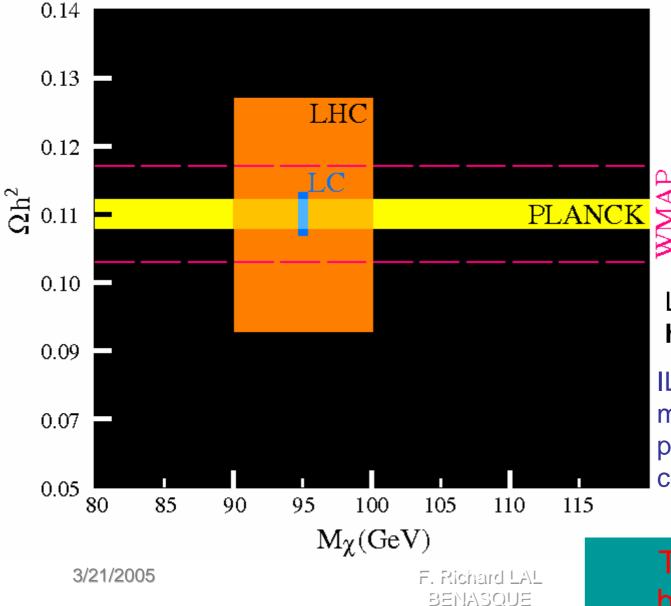
adam magy (int)

Direct: DAMA



F. Richard LAL Nuclear Spin dependence ?

Colliders and Cosmology



MicrOMEGAs Pt B

'WMAP'	7 %
LHC	~15 %
'Planck'	~2 %
CL	~3 %

LHC pt B: Battaglia et al hep-ph/0306219

ILC: precision similar for most other co-annihil. points A C D G I L consistent with WMAP

Theory: could be a limiting factor

Complementarities

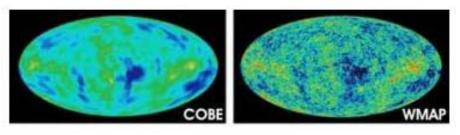
- The DM example illustrates how ILC can contribute to the DM issue in cosmology
- Similarly, If SUSY is the correct theory, it can also:
 - Explain Baryogenesis provided that there are light \widetilde{t}_1 +chargino and extra phases
 - Provide a mechanism to give masses to neutrinos

In both cases one needs

the ILC to investigate the scenario

www.

Precision measurements @ILC



- e.g. (MH)Direct≠(MH)Indirect±5 GeV ?
 -> New physics is present
- Which type of new physics ?
- Which new mass scale ?
- ILC provides several precise observables to provide an answer
- Take an example: a Z'

The Z' scenario

- Expected in a large class -0.25 of models, including SUSY (origin of μ in $\mu H_u H_d$ pb)
- (origin of μ in $\mu H_u H_d$ pb) • Indirect detection with high precision ILC allows to go to ~10 TeV and also

0.5

0.25

0

-0.4

-0.2

LR

m_{z'} = 1.5 TeV

0.4

02

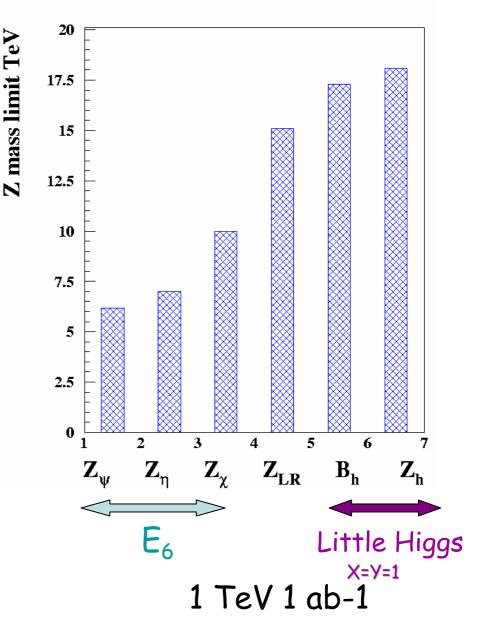
0 a'ı

- ILC essential to identify the type of Z'
- -> Define future machines (CLIC, VLHC)

LHC :

- up to ~5 TeV direct observation

- up to ~2 TeV identif.LC :
- discriminate between
 models up to ≥ 5 TeV
- predict MZ' with a relative accuracy
- < (MZ'/10TeV)²
- < 25 % at 5 TeV



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Possible scenarios at LHC with the consequences on TeV LC called TLC (and on CLIC)

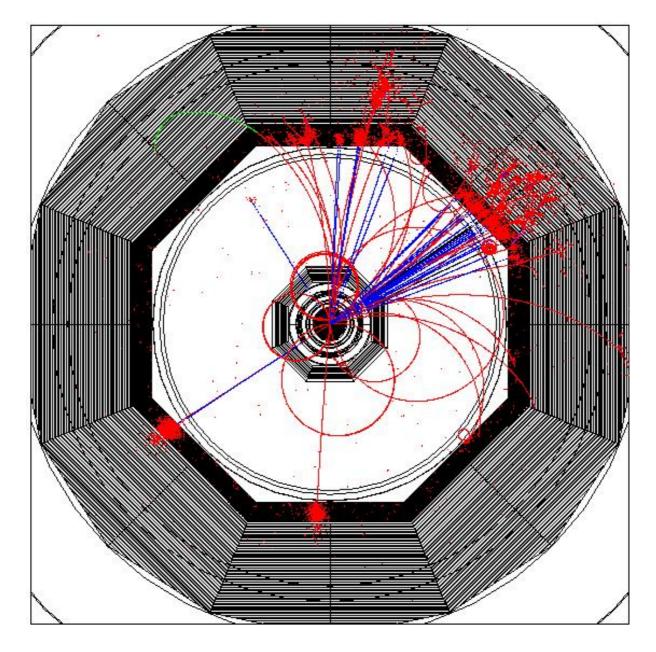
Scenario	Models	GigaZ at TLC	TLC below 500 GeV	High Energy TLC	LHC	From LHC to CLIC (or SLHC)
Light Higgs alone		Quantum Test (QT) ∆M _H =5 GeV > SM ?	Spin/Parity BR(fermions.g/y,W*W) ZZH and WWH couplings	ZHH 10% ttH <10% anomalous W/Z couplings	~10 times less precise, less BR's no spin/parity ~ no ZHH	Not motivated (SLHC for ZHH)
Light Higgs +SUSY		Stop mass,A _t Full info available	Indirect mA, mH Same except WWH	Same+A,H,H [*] Heavy Higgses (e.g. into γ [*])	Same Difficult, no Higgs mass information	Heavier Higgses Heavier Higges
		QT QT If mH< 300 GeV: Room for Physics beyond SM ?	ZZH reduced ZZH, WWH up to 400 GeV	Heavy Higgses Higgs up to ~0.8 TeV anomalous W/Z couplings	More Difficult Confirm SM but loweraccuracy no QT info	Heavier Higgses Not Motivated
Heavy Higgs	(SUSY) Little Higgs		Idem Idem	Same+additional Higgses A.H.H* Z' up to 15 TeV	Ambiguous unless A,H,H [±] identified Z ⁺ up to ~5 TeV	Not motivated if
		to 7 TeV QT Z-Z' mix for Z' up to 13 TeV	Idem	Effect on ZWW Z' up to 20 TeV	Ambiguous Z' up to ~5 TeV Ambiguous	Z' not seen at LHC (SLHC motivated Z') Not motivated if Z' not seen at LHC (as above for Z')
	ND>4 UED	ρ effect isospin violation for KK masses <1 TeV	Idem	Pair production of KK excitations		CLIC extends LHC/LC mass range
No Higgs	- Higgless	QT -> Constraints on theory > SM QT -> Constraints on theory	WLWL >3 σ larger with resonance	WLWL >6& larger with resonance Z' visible ZWW anomaly	>5σ Observation of Z' but ambiguous	Direct observation of p-resonance(?) Z' on shell at CLIC

Detector

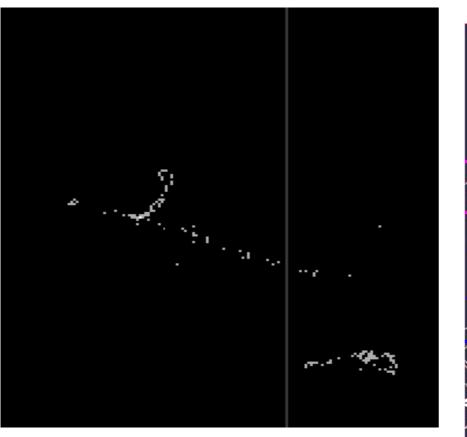
In many instances LC analyses will be systematics limited

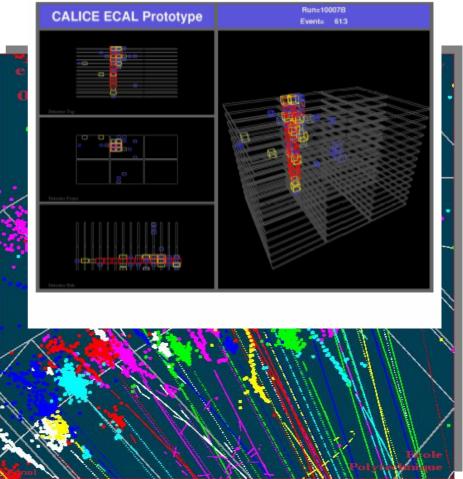
3 outstanding improvements/LEP-SLD can be fulfilled on the detectors:

- Improved vertexing : c (ε=70% >80% pure), tau tagging
- Improved E-Flow : 6/8 jets WW/ZZvv δE/E~0.3/√E
- δp/p² ~ 1/10 LEP down to 100 mrad
 Also:
- Hermeticity on energetic γ /e down to 5 mrad
- AL/L to 10⁻⁴ + Polar + vs to 1 MeV for Z physics
 -> Machine Detector interface



Towards a bubble chamber pattern ?

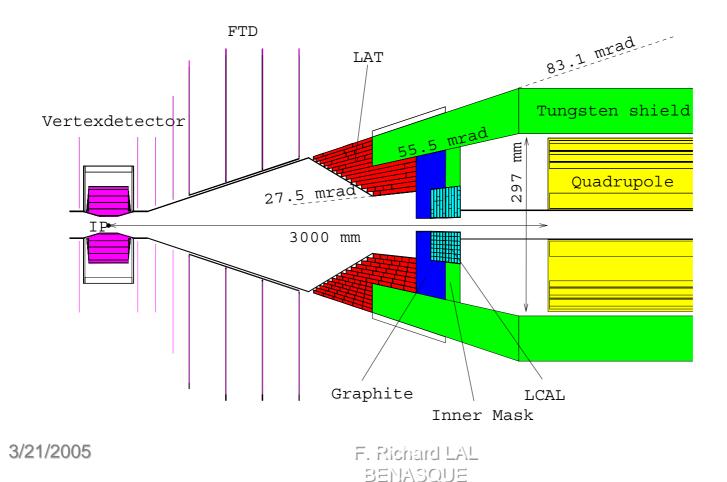




Pixelized TPC 3/21/2005

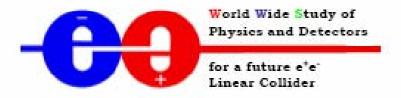
Calorimetry with high granularity

Forward Hermeticity down to 5 mrad

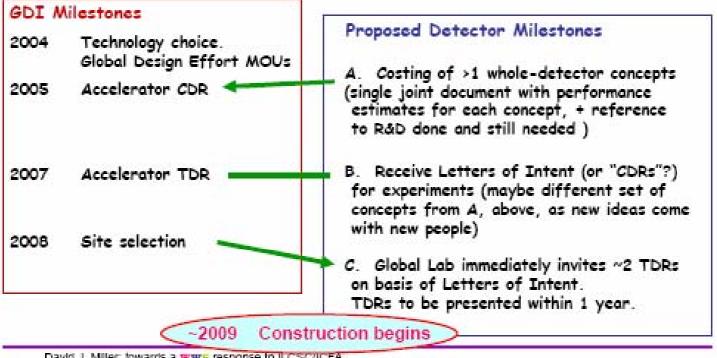


What is happening in the world?

- 3 concepts of detectors are discussed, one is ~TESLA TDR, one with a Si tracker in the US, one considering a larger detector
- An international R&D panel is been set up to identify the priorities
- A CDR for the machine and a document for the detectors beginning of 2006
- Lol in ~2008 sent to the Global Lab
- 2 experimental areas are envisaged



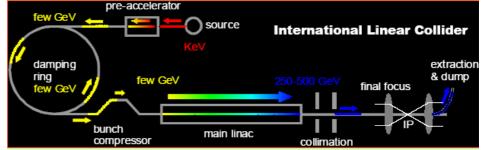
We propose to tie detector milestones to the Global LC Design Initiative.



David J. Miller; towards a wws response to ILCSC/ICFA

Critical items for the machine

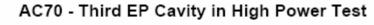
- Major investment (>80%) is on the SC Linac
- There are however many other issues e.g.
 - Polarized e-
 - Damping Rings
 - Generation of e+
 - Beam Delivery
 -

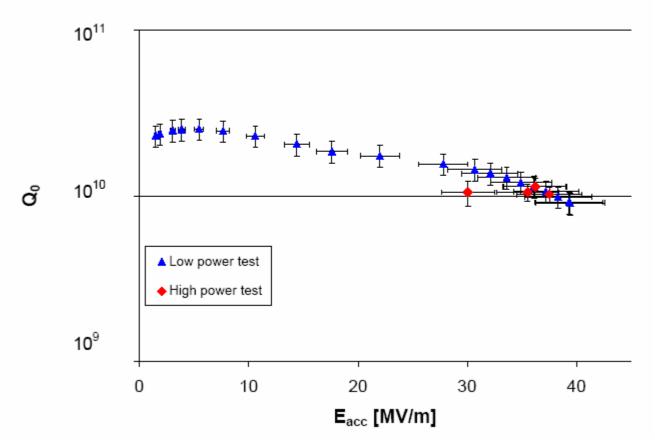


- The recommendation on technology allows to share the work among the 3 regions
- New evaluation after the TESLA TDR but one should avoid 'reinventing the wheel'
- Needs leadership, coordination, identification of resources, sharing of the work -> GDE (global design effort°

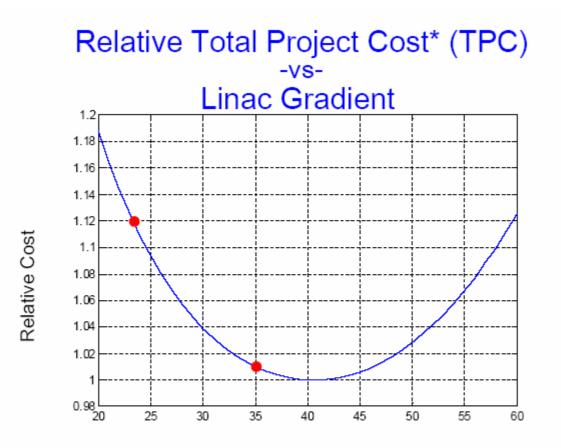


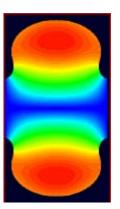












Gradient (MV/m) •TPC is for 500 GeV machine in US Options Study but does not include additional unpowered tunnel sections.

How to build a TeV LC ?

- Perform active R&D during ~10 Years
- Go international (\cost, ↗effort) ICFA->ILCSC
- Create a large community of users -> ECFA in Europe + LCWS workshops
- Decide that the LC is the highest priority of HEP
- -> Conclusion from ECFA after a wide consultation
- Get international recognition OCDE, EU
- Choose the best technology ITRP August 2004
- Join efforts to construct the best machine: Recent meeting at KeK
- Convince the scientific community, the public, the politicians -> outreach
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 F. Richard LAL



- Consultations ECFA organised in 2001 on the request of CERN and DESY directors
- Document ECFA (endorsed by the Consultative group OCDE and acknowledged by Ministers of Science in Jan 2004):

	M
() i)	Exploiting the current frontier facilities until the contribution from these machines is surpassed by the results from the LHC or LC;
ii)	Completing and then fully exploiting the LHC;
iii)	Preparing for the approval of a Linear Collider of at least 400 GeV centre of mass energy, to run concurrently with the LHC in the decade starting in 2010;
iv)	Supporting an appropriate R&D programme into novel accelerator designs (for very high energy electron-positron linear colliders, neutrino factories, muon colliders and very high energy hadron colliders).
3/21/2005	F. Richard LAL BENASQUE



A Community for the ILC

Euro Collaborations

TESLA (wider than Europe alone)

European XFEL

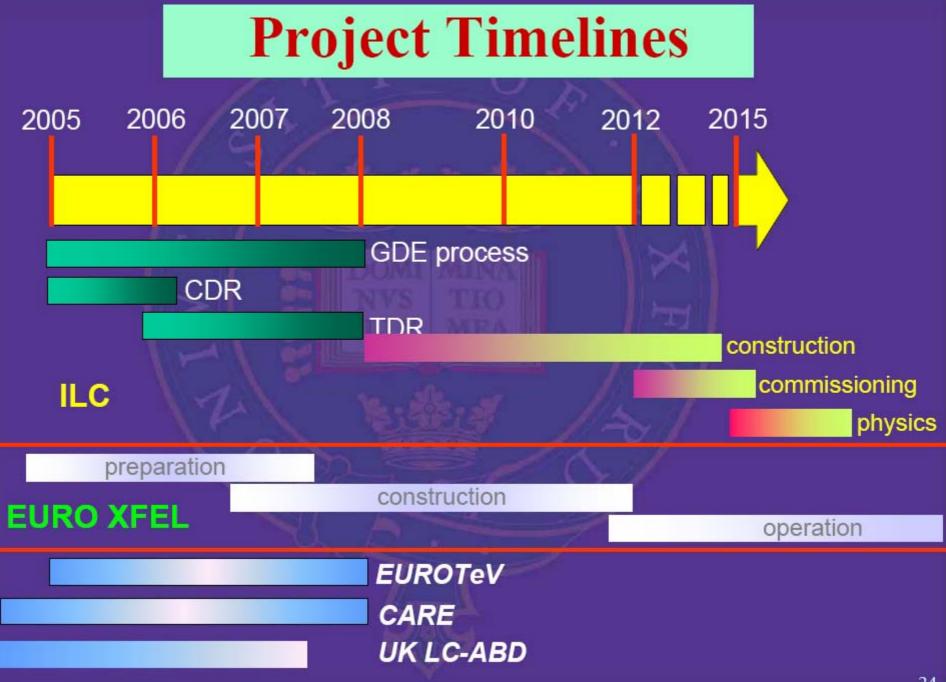
Coordinated Accelerator Research in Europe

• EuroTeV - LC research programme

UK Linear Collider Accelerator & Beam Delivery LCABD – PPARC & CCLRC-funded

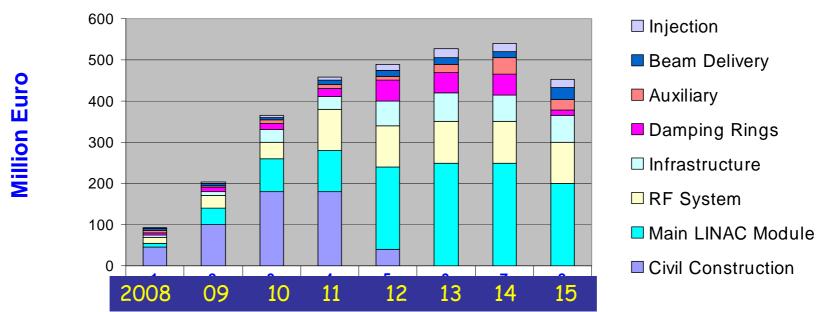






A. Wagner CB TESLA April 2004

An exercise.....



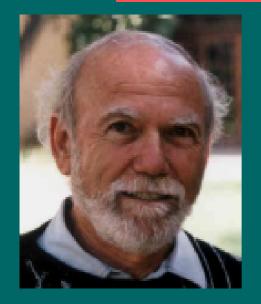
TESLA material cost vs construction year

This is assuming a construction time of 8 years.

By parallel manufacturing of components this construction time can be shortened to ~ 6 years

-> matches turn on and first results of LHC before major spending starts

Our prioritization of the candidates was:



1. Barry Barish

2. Satoshi Ozaki





3. David Burke

Conclusions

- Physics arguments for a TeV ILC at very high luminosity are based on the Higgs sector
- There is no guarantee that new physics is directly observable at LHC/ILC
- LC can provide essential inputs, through PM, for indirect observations on new physics (Z')
- If light SUSY, ILC essential for accuracy (DM)
- The international process is continuously progressing and Europe has the SC technology (TTF, XFEL) well in hand and resources from EU
- I hope that Spain will become a major player in this fascinating enterprise

International Recognition for WW LC

Science, Technology and Innovation for the 21st Century Meeting of the OECD Committee for Scientific and Technological Policy at Ministerial Level, 29-30 January 2004 – Final Communique

They noted the worldwide consensus of the scientific community, which has chosen an electron-positron linear collider as the next acceleratorbased facility to complement and expand on the discoveries that are likely to emerge from the Large Hadron Collider currently being built at CERN. They agreed that the planning and implementation of such a large, multiyear project should be carried out on a global basis, and should involve consultations among not just scientists, but also representatives of science funding agencies from interested countries. Accordingly, Ministers endorsed the statement prepared by the OECD Global Science Forum Consultative Group on High-Energy Physics:

A **roadmap** that identifies four interdependent priorities for global highenergy physics (HEP) facilities:

-The exploitation of current frontier facilities until contribution of these machines is surpassed

- Completion and full exploitation of the Large Hadron Collider at CERN

- Preparing for the development of a **next-generation electron-positron** collider

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M. Tigner LP2003

R1: R&D Needed for a Feasibility Demonstration of the Machine

R1 'Score Card': Is a Feasibility Demonstration Required*?

	Modulators	Klystrons	RF Distribution	Accelerator Structures
TESLA	No	No	No	No (500 GeV) Yes (800 GeV)
NLC/JLC-X	No	No	Yes	Yes
JLC-C	No	No	Yes	Yes
CLIC	Yes	Yes	Yes	Yes

	TABLE 2: Sumn				
	\mathbf{TESLA}		$ m JLC-X/NLC^a$		
Center of mass energy [GeV]	500	800	500	1000	
RF frequency of main linac [GHz])	1.3	3	11	11.4	
Design luminosity $[10^{33} \text{ cm}^{-2} \text{s}^{-1}]$	34.0	58.0	25.0(20.0)	25.0(30.0)	
Linac repetition rate [Hz]	5	4	150(120)	100(120)	
Number of particles/bunch at IP $[10^{10}]$	2	1.4	0.	75	
$\gamma \varepsilon_x^* / \gamma \varepsilon_y^*$ emit. at IP [m·rad × 10 ⁻⁶]	10 / 0.03	8 / 0.015	3.6 /	0.04	
$\beta_x^{\star} / \beta_y^{*}$ at IP [mm]	15 / 0.40	15 / 0.40	$\{ 8 / 0.11 \}$	13 / 0.11	
$\sigma^{\star}_{\boldsymbol{x}} / \sigma^{*}_{\boldsymbol{y}}$ at IP [nm] before pinch ^c	554 / 5.0	392 / 2.8	$2\ 243\ /\ 3.0$	219 / 2.1	
σ_z^{\star} at IP [μ m]	300)	1	10	
Number of bunches/pulse	2820	4886	1	92	
Bunch separation [nsec]	337	176	1	.4	
Bunch train length $[\mu sec]$	950	860	0.2	267	
Beam power/beam [MW]	11.3	17.5	8.7~(6.9)	11.5(13.8)	
Unloaded/loaded gradient ^{d} [MV/m]	23.8 / 23.8 ^e	35 / 35	4 65	/ 50	
Total number of klystrons	572	1212	4064	8256	
Number of sections	20592	21816	12192	24768	
Total two-linac length [km]	30	30	13.8	27.6	
Total beam delivery length [km]	3		3	.7	
Proposed site length [km]	33		3	2	
Total site AC power ^f [MW]	140	200	243 (195)	292 (350)	
Tunnel configuration ^g	Sing	le	Do	uble	

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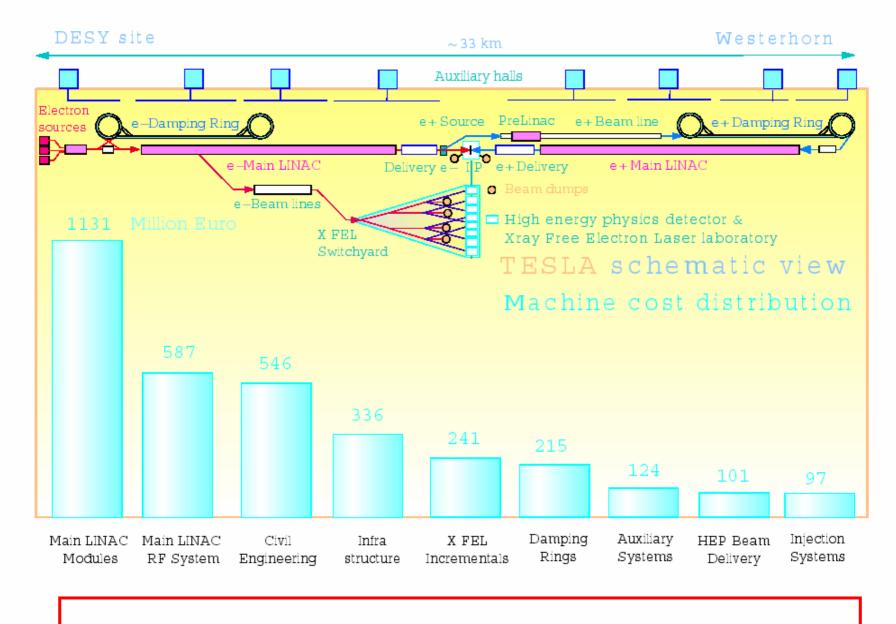


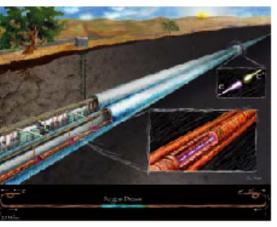
Figure 4.1.2:

Overview of the accelerator investment costs

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Mid-Term Priorities

Priority: 13 Linear Collider



The Linear Collider is designed to extend the study of particle physics.

1st priority of Mid-Term

The Facility: The Linear Collider will allow physicists to make the world's most precise measurements of nature's most fundamental particles and forces at energies comparable to those of the Large Hadron Collider (LHC) now under construction in Switzerland.

Background: The Standard Model of particle physics, developed over the last 50 years and recognized as one of *the* great scientific achievements, has been tremendously effective in predicting the behavior of all the interactions of subatomic particles except those due to gravity, and in describing the varieties of particles that combine to make everyday matter. The next step—incorporating a theory of gravity and understanding why fundamental particles have mass—will require particle accelerators that function at the trillion-electron volt ("TeV") level.

Facilities for the Future of Science A Twenty-Year Outlook

1			
Î	Category A Highest Scientific Importance Near-term Readiness for Construction	Category B Highest Scientific Importance Mid-term Readiness for Construction	Category C Highest Scientific Importance Far-term Readiness for Construction
of Science	◀	Category D Secondary Scientific Importance Varying Readiness for Construction	
Importance	◀	Category E Hard to Assess Scientific Importance Varying Readiness for Construction	
	Readin	ess for Construction -	→

Office of Science Facilities Matrix

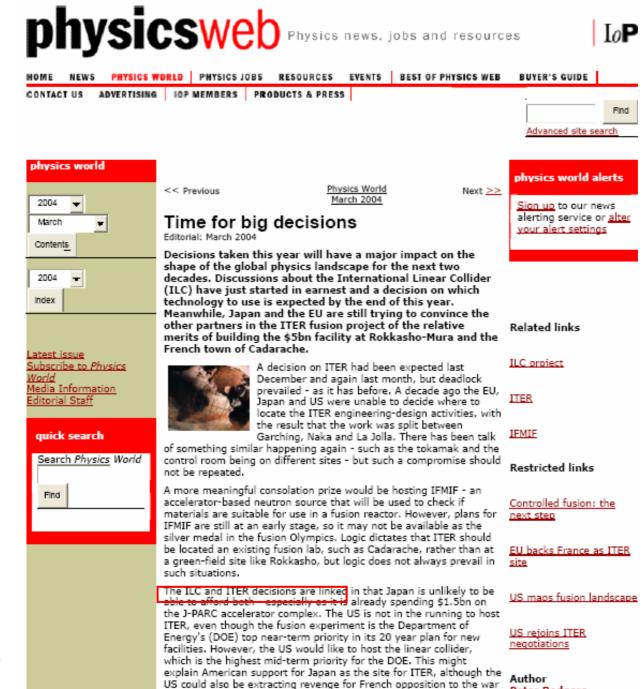






November 2003

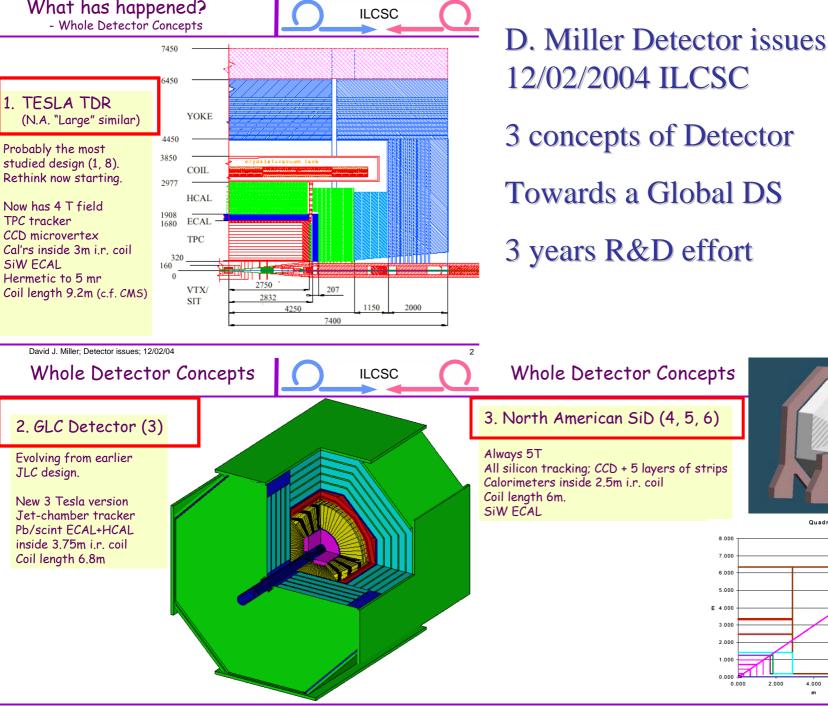
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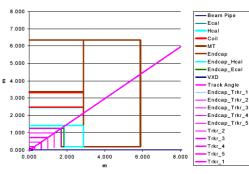
in Iraq.

Peter Rodaers

3/21/2005



Quadrant View



3

R&D programmes



(taken from WWS R&D website. http://blueox.uoregon.edu/~lc/randd.html list incomplete - due for updates)

WWS keeping them in touch, encouraging co-operation between regions.

VERTEX DETECTOR

KEK-led collaboration

LCFI collaboration UK

RAL-led collaboration

European collaboration

Monolithic APS

Bonn/MPI Group

Hybrid APS

DEPFET

Oregon/Yale Collaboration

Strasbourg-led collaboration

CCD

MAIN TRACKER

TPC

<u>Aachen, LBNL, MIT, DESY/U.Hamburg</u> <u>Carleton/Montreal/Victoria</u>, <u>CERN</u> <u>Orsay</u>, <u>Saclay</u>, <u>Wayne State</u>, <u>MPI-Munich</u>, <u>Japan</u>

Silicon

<u>LPNE Paris</u>, Santa Cruz, UCSC, Michigan, Wayne State, SLAC, Asian groups

Jet Chamber Asia

David J. Miller; Detector issues; 12/02/04





CALORIMETRY

SiW ECAL (+ HCAL) CALICE, 28 Labs from 8 countries, including Europe, US, Canada and Korea. SiD, North America

Tiles etc. Padova KEK et al (GLC) ALSO FORWARD DETECTORS MUON DETECTORS PARTICLE IDENTIFICATION TRIGGER+DATA ACQUISITION TEST BEAMS GAMMA GAMMA DETECTOR BEAMLINE INSTRUMENTATION

MOST R&D programmes are underfunded.

Not enough test beams available, especially with high energy hadrons.

David J. Miller; Detector issues; 12/02/04

Z' at LHC/ILC

- Below 1 TeV gain for low coupling by Rad Return
- Above 5 TeV gain in mass with PM

